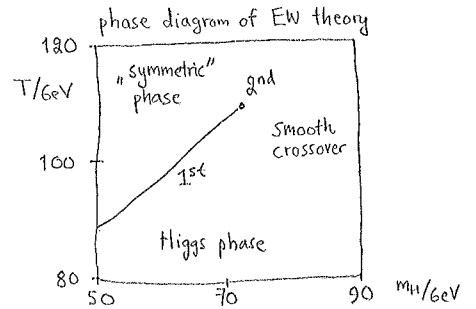
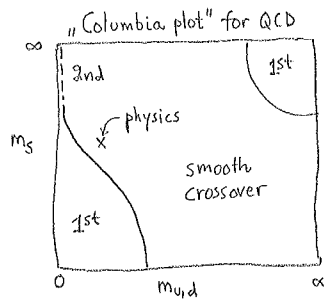


Thermal phase transitions. 1. Theoretical tools

Overall picture:

Both QCD and the electroweak (EW) theory display a "broken symmetry" in vacuum, which gets "smoothly restored" at high temperatures:



In extensions of the electroweak theory, things might go otherwise. Given a model, we may ask:

- (i) is there an actual phase transition? critical temperature
- (ii) what are its "equilibrium properties" ( $T_c, L, \xi$ )? latent heat
- (iii) how does it proceed in real time? surface tension
- (iv) does it leave remnants (gravitational waves, baryon asymmetry)?

Quantum statistical physics: Let us define and compute a number of basic objects:

- \* partition function:  $Z = \text{Tr}(e^{-\beta \hat{H}})$ ,  $\beta := \frac{1}{T}$ ,  $\hat{H}$  = Hamiltonian
- \* free energy:  $Z = e^{-\beta F} \Leftrightarrow F = -T \ln Z$
- \* free energy density:  $f = \frac{F}{V}$ ,  $V = \text{volume} \rightarrow \infty$

We compute these for a harmonic oscillator:

$$\hat{H} = \frac{\hat{p}^2}{2m} + \frac{m\omega^2 \hat{x}^2}{2} = \hbar\omega \left( \hat{a}^\dagger \hat{a} + \frac{1}{2} \right)$$

Denote  $\epsilon := \hbar\omega$ . Then

$$Z = \sum_{n=0}^{\infty} \langle n | e^{-\beta \hat{H}} | n \rangle = \sum_{n=0}^{\infty} e^{-\beta \epsilon \left( n + \frac{1}{2} \right)} = \frac{e^{-\beta \epsilon / 2}}{1 - e^{-\beta \epsilon}}$$

$$F = \frac{\epsilon}{2} + T \ln(1 - e^{-\beta \epsilon})$$

We also need the analogue of a "propagator". In order to be able to use this later on in field theory, let us set  $m, \hbar \rightarrow 1$ .

Then the coefficient of  $\hat{x}^2$  in  $\hat{H}$  is  $\frac{\epsilon^2}{2}$ , and

$$\begin{aligned} \langle \hat{x}^2 \rangle &:= \frac{\text{Tr}(\hat{x}^2 e^{-\beta \hat{H}})}{Z} = -2T \frac{\partial}{\partial \epsilon^2} \ln Z \\ &= \frac{1}{\epsilon} \frac{\partial F}{\partial \epsilon} = \frac{1}{\epsilon} \left[ \frac{1}{2} + \frac{e^{-\beta \epsilon}}{1 - e^{-\beta \epsilon}} \right] \\ &= \frac{1}{\epsilon} \left[ \frac{1}{2} + \frac{1}{e^{\beta \epsilon} - 1} \right] \\ &= n_B(\epsilon) \quad (\text{Bose distribution}) \end{aligned}$$

Statistical field theory: We can take a short cut from the harmonic oscillator to (free) scalar field theory, simply by replacing  $\epsilon \rightarrow \epsilon_k := \sqrt{k^2 + m^2}$ .

$$Z_\phi = \int \frac{d\phi}{\mathcal{R}} \exp \left\{ -\frac{1}{T} \left[ \frac{\epsilon_k}{2} + T \ln(1 - e^{-\beta \epsilon_k}) \right] \right\},$$

$$f_\phi = \lim_{V \rightarrow \infty} \frac{F_\phi}{V} = \lim_{V \rightarrow \infty} \frac{1}{V} \sum_k \underbrace{\left[ \frac{\epsilon_k}{2} + T \ln(1 - e^{-\beta \epsilon_k}) \right]}_{\sim E - TS}$$

\* L. Dolan and R. Jackiw, "Symmetry behavior at finite temperature", Phys. Rev. D 9 (1974) 3320

The thermal part  $f_T(m) := T \int_k \ln(1 - e^{-\beta \epsilon_k})$  is exponentially convergent. It has a formal expansion\* which is non-analytic in  $m^2$ :

$$f_T(m) = -\frac{\pi^2 T^4}{90} + \frac{m^2 T^2}{24} - \frac{m^4 T}{12\pi} - \frac{m^4}{32\pi^2} \left[ \ln\left(\frac{m}{4\pi T}\right) + \gamma_E - \frac{3}{4} \right] + \mathcal{O}\left(\frac{m^6}{\lambda^4 T^2}\right)$$

*m := sqrt(m^2)*      *like -p on p. 2*

Example: Higgs condensate

We may use the above expansion to estimate  $\langle \phi^\dagger \phi \rangle$  at high T:

$$\langle \phi^\dagger \phi \rangle \approx \frac{1}{2} \sum_{n=0}^3 \int_k \frac{1}{\epsilon_k} \left[ \frac{1}{2} + n_B(\epsilon_k) \right]$$

$\phi = \frac{1}{\sqrt{2}} (\phi_0 + i\phi_3)$       *p. 29:  $2 \frac{\partial}{\partial \epsilon_k^2} \left[ \frac{\epsilon_k}{2} + T \ln(1 - e^{-\beta \epsilon_k}) \right]$ ;  $\frac{\partial}{\partial \epsilon_k^2} = \frac{\partial}{\partial m^2}$*

$$= (\text{vacuum part}) + \sum_{n=0}^3 \frac{\partial}{\partial m^2} f_T(m) = \frac{T^2}{6} + \mathcal{O}(mT)$$

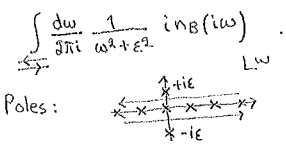
Therefore, at  $T \gg v \approx 246$  GeV, zero-temperature masses  $m_W^2 = \frac{g^2 v^2}{4}$  may get replaced with thermal masses  $\sim g^2 T^2$  (more later).

Issues of convergence:

In perturbation theory, the most important domain is where propagators are largest:

$$\frac{1}{\epsilon_k} \left[ \frac{1}{2} + \frac{1}{e^{\epsilon_k/T} - 1} \right] \approx \frac{1}{\epsilon_k} \left[ \frac{1}{2} + \frac{1}{\epsilon_k/T + \epsilon_k^2/24T^2 + \dots} \right] = \frac{T}{\epsilon_k^2} + \mathcal{O}\left(\frac{1}{T}\right)$$

\*\* This can be verified by considering the contour integral

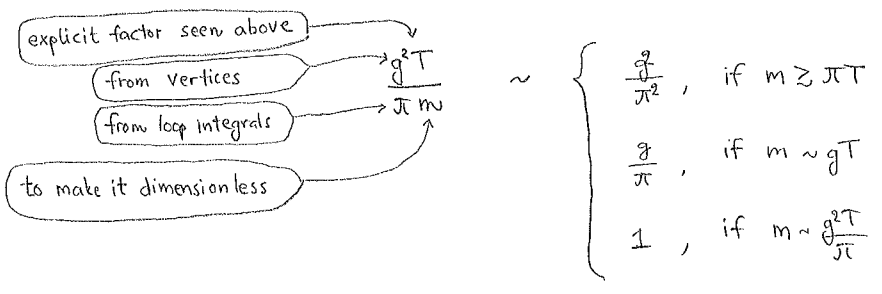


Closing the contours in two different ways yields the desired equality.

The large term at  $\epsilon_k \ll T$  originates from Bose-Einstein-enhancement. But there is also a very useful alternative interpretation\*\*:

$$\frac{1}{\epsilon_k} \left[ \frac{1}{2} + \frac{1}{e^{\epsilon_k/T} - 1} \right] = T \sum_{\omega_n} \frac{1}{\omega_n^2 + \epsilon_k^2}, \quad \omega_n = 2\pi n T, \quad n \in \mathbb{Z}$$

Then the large term is associated with the "Matsubara zero mode"  $\omega_n = 0$ . With these ingredients, we can estimate a dimensionless loop expansion parameter (after integration over momenta, so that  $m$  is the only scale):



\*\*\* A.D. Linde, "Infrared problem in thermodynamics of Yang-Mills gas", Phys. Lett. B 96 (1980) 289

The breakdown of the loop expansion for very small masses is known as the "Linde problem". \*\*\*

Imaginary-time formalism:

It can be shown (replacing  $e^{-\frac{i\hat{H}t}{\hbar}} \rightarrow e^{-\beta\hat{H}}$  in the usual derivation) that the partition function can be expressed as a "Euclidean"/"imaginary-time" path integral:

$$Z = \int_{b.c.} \mathcal{D}\phi e^{-S_E}$$

$$S_E = \int_0^\beta d\tau \int_V d^3x L_E$$

$$L_E := -\mathcal{L}_M(it \rightarrow \tau)$$

Here boundary conditions (b.c.) are periodic for bosons and antiperiodic for fermions. Check:

$$\phi(\beta, \vec{x}) = \phi(0, \vec{x}) \Rightarrow e^{i\omega_n \beta} = 1 \Rightarrow \omega_n = 2\pi T n \quad \Delta k!$$

We apply this to a real scalar field with  $V(\phi) = -\frac{m^2 \phi^2}{2} + \frac{\lambda \phi^4}{4}$ :

$$\mathcal{L}_M = \frac{1}{2} (\partial_\mu \phi)^2 - \frac{1}{2} (m\phi)^2 - V(\phi)$$

$$\Rightarrow L_E = \frac{1}{2} (\partial_\tau \phi)^2 + \frac{1}{2} (\nabla \phi)^2 + V(\phi)$$

Effective potential:

We return to a finite volume for a moment. Let  $\bar{\phi}$  be the mode with  $\omega_n=0, \vec{k}=\vec{0}$ , and  $\phi'$  the other modes. Note that  $\int_0^\beta d\tau \int_V d^3x \phi' = 0$ ! Now insert  $\phi = \bar{\phi} + \phi'$  into  $Z$ :

$$Z = \int_{-\infty}^{\infty} d\bar{\phi} \int \mathcal{D}\phi' e^{-S_E[\phi = \bar{\phi} + \phi']} =: \int_{-\infty}^{\infty} d\bar{\phi} e^{-\frac{V}{T} V_{eff}(\bar{\phi})}$$

$$=: \int_{-\infty}^{\infty} d\bar{\phi} e^{-\frac{V}{T} [V_{eff}(\bar{\phi}_{min}) + \frac{1}{2} V_{eff}''(\bar{\phi}_{min})(\bar{\phi} - \bar{\phi}_{min})^2 + \dots]}$$

$$\Rightarrow f_\phi = V_{eff}(\bar{\phi}_{min}) + \mathcal{O}\left(\frac{\hbar N V}{V}\right)$$

Insert  $\phi = \bar{\phi} + \phi'$  in  $L_E$ :

$$\begin{aligned} \frac{1}{2} (\partial_\mu \phi)^2 &\rightarrow \frac{1}{2} (\partial_\mu \phi')^2 \\ -\frac{1}{2} m^2 \phi^2 &\rightarrow -\frac{1}{2} m^2 \bar{\phi}^2 - m^2 \bar{\phi} \phi' - \frac{1}{2} m^2 \phi'^2 \\ +\frac{1}{4} \lambda \phi^4 &\rightarrow +\frac{1}{4} \lambda \bar{\phi}^4 + \lambda \bar{\phi}^3 \phi' + \frac{3}{2} \lambda \bar{\phi}^2 \phi'^2 + \lambda \bar{\phi} \phi'^3 + \frac{1}{4} \lambda \phi'^4 \end{aligned}$$

independent of  $\tau, \vec{x}$ 
vanishes because  $\int d\tau \int d^3x \phi' = 0$ 
like free theory but with mass  $m_{eff}^2 := -m^2 + 3\lambda \bar{\phi}^2$ 
interactions

In summary, the effective potential  $V_{eff}$  is composed of the following parts:

\* tree-level potential:  $V_{eff}^{(0)}(\bar{\phi}) = -\frac{1}{2} m^2 \bar{\phi}^2 + \frac{1}{4} \lambda \bar{\phi}^4$

\* 1-loop potential: like on p.30 but now with  $m_{eff}$ :

$$V_{eff}^{(1)}(\bar{\phi}) = \int_{\vec{k}} \left[ \frac{\epsilon_k}{2} + T \ln(1 - e^{-\beta \epsilon_k}) \right]$$

$$\epsilon_k = \sqrt{k^2 + m_{eff}^2}$$

\* higher order corrections from the interaction part.

Phase transition:

Let us insert the high-temperature expansion from p.30 and see what kind of an effect  $V_{\text{eff}}^{(1)}$  has:

\* leading term  $-\frac{\pi^2 T^4}{90} \Rightarrow$  independent of  $\bar{\phi} \Rightarrow$  no effect

\* NLO term  $\frac{m_{\text{eff}}^2 T^2}{24} = \frac{-m^2 + 3\lambda \bar{\phi}^2}{24} T^2$

$$\Rightarrow V_{\text{eff}}^{(0)} + V_{\text{eff}}^{(1)} \approx [\bar{\phi}\text{-indep.}] + \frac{1}{2} \left( -m^2 + \frac{\lambda T^2}{4} \right) \bar{\phi}^2 + \frac{1}{4} \lambda \bar{\phi}^4$$

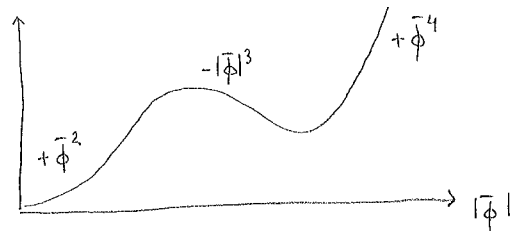
↑  
"thermal mass correction"

Thus the symmetry seems to get restored at  $T_c \approx \frac{2m}{\sqrt{\lambda}}$ .

\* NNLO term  $-\frac{m_{\text{eff}}^3 T}{192\pi} = -\frac{T}{192\pi} (-m^2 + 3\lambda \bar{\phi}^2)^{3/2}$

Let us for simplicity put  $m^2 \rightarrow 0$

$$\Rightarrow V_{\text{eff}}^{(0)} + V_{\text{eff}}^{(1)} \approx [\bar{\phi}\text{-indep.}] + \frac{\lambda}{8} T^2 \bar{\phi}^2 - \frac{T}{192\pi} (3\lambda)^{3/2} |\bar{\phi}|^3 + \frac{\lambda}{4} \bar{\phi}^4$$



This looks a bit like a "fluctuation induced" / "radiatively generated" first order phase transition!

However at the end we have to ask whether such a prediction is reliable? Rough estimate:

\* "broken minimum" is where  $\frac{T}{\pi} \lambda^{3/2} |\bar{\phi}|^3 \sim \lambda \bar{\phi}^4$

$$\Rightarrow |\bar{\phi}| \sim \lambda^{1/2} \frac{T}{\pi}$$

\* dimensionless loop expansion parameter (p.30) is

$$\frac{\lambda T}{\pi m_{\text{eff}}} \sim \frac{\lambda T}{\pi \sqrt{\lambda \bar{\phi}^2}} \sim \frac{\lambda^{1/2} T}{\pi |\bar{\phi}|} \sim 1$$

So we conclude that the prediction for a first-order transition is not reliable (series does not converge).

In fact, in scalar field theory the transition is of 2<sup>nd</sup> order (continuous).