

### 8.1 A sample computation

Let us now apply the relations found to a practical computation. Two possible formulations:

(i) "Imaginary time followed by analytic continuation"

⇒ Start by computing the "easy" observables  $\Pi_E(i\omega) / \tilde{\Pi}_E(\tau)$  in the imaginary time (Matsubara) formalism. Then apply the relations obtained to this result, to get  $S(p)$ .

(ii) "Real-time formalism"

⇒ Formulate the analytic continuation on the level of Feynman rules already. Naively this means carrying out computations as at zero temperature in Minkowskian spacetime, but with a thermal correction in the propagator (p.108):

$$-i\Pi_T(p_0) = \frac{+i}{p_0^2 - E^2 + i\epsilon} + 2\pi \delta(p_0 - E^2) n_b(|p_0|)$$

↓ as at T=0
↑ temperature

Unfortunately a more careful analysis reveals that this shortcut does not work in general: one needs to introduce several propagators, corresponding to  $\Pi^>, \Pi^<, \dots$ .

[see the book by LeBellac, for instance]

People tend to have (strong) personal preferences as to which way is better. In the following we use (i).

Let us now consider the following example: [H.A. Weldon, Phys. Rev. D 28 (1983) 2007]

$\mathcal{L}_I \sim g\phi\chi_1\chi_2$

$\phi \xrightarrow{x_1} \text{---} \text{---} \text{---} \phi$   
 $\text{---} \text{---} \text{---} \chi_2 \rightarrow Q = (q_n, \bar{q})$

$$T_{E}(q_n, \bar{q}) \equiv g^2 T \sum_{p_n} \int \frac{d^3 \vec{p}}{(2\pi)^3} \frac{1}{[p_n^2 + p_0^2][(p_n + q_n)^2 + r_0^2]} ; \quad \begin{aligned} p_0^2 &\equiv E_1^2 \equiv m_1^2 + \vec{p}^2 \\ r_0^2 &\equiv E_2^2 \equiv m_2^2 + (\vec{p} + \vec{q})^2 \end{aligned}$$

(1) Carry out  $T \sum_{p_n}$ , using

$$T \sum_{p_n} f(p_n) = \frac{1}{2} \int_{-\infty - i\epsilon}^{\infty - i\epsilon} \frac{dw}{2\pi i} [f(w) + f(-w)] [1 + 2n_L(iw)]$$

Factor out  $g^2 \int \frac{d^3 \vec{p}}{(2\pi)^3}$  for a moment.

Write:

$$f(w) + f(-w) = \frac{1}{(w^2 + p_0^2)} \left[ \frac{1}{(w + q_n)^2 + r_0^2} + \frac{1}{(-w + q_n)^2 + r_0^2} \right]$$

$$= \frac{1}{(2p_0)(2r_0)} \left[ \frac{i}{w + ip_0} - \frac{i}{w - ip_0} \right] \times \left[ \frac{i}{w + q_n + ir_0} - \frac{i}{w + q_n - ir_0} + \frac{i}{-w + q_n + ir_0} - \frac{i}{-w + q_n - ir_0} \right]$$

Pick up poles in the lower half plane:

$$\frac{1}{2} \cdot \frac{1}{2\pi i} \cdot (-2\pi i) \cdot \frac{i}{(2p_0)(2r_0)} \left\{ [w = -ip_0] (1 + 2n_L(p_0)) \left( \frac{i}{q_n - ip_0 + ir_0} - \frac{i}{q_n - ip_0 - ir_0} + \frac{i}{q_n + ip_0 + ir_0} - \frac{i}{q_n + ip_0 - ir_0} \right) \right.$$

Note:  $n_L(\pm q_n + r_0) = n_L(r_0)$  !!

$$[w = -q_n - ir_0] (1 + 2n_L(r_0)) \left( -\frac{i}{q_n - ip_0 + ir_0} + \frac{i}{q_n + ip_0 + ir_0} \right)$$

$$[w = q_n - ir_0] (1 + 2n_L(r_0)) \left( \frac{i}{q_n + ip_0 - ir_0} - \frac{i}{q_n - ip_0 - ir_0} \right)$$



$$= \frac{1}{(z_p)(z_{r_0})} \left\{ \frac{1}{-iq_n - p_0 + r_0} \left[ \frac{1}{2} + n_b(p_0) - \frac{1}{2} - n_b(r_0) \right] \right. \\ + \frac{1}{iq_n + p_0 + r_0} \left[ \frac{1}{2} + n_b(p_0) + \frac{1}{2} + n_b(r_0) \right] \\ + \frac{1}{-iq_n + p_0 + r_0} \left[ \frac{1}{2} + n_b(p_0) + \frac{1}{2} + n_b(r_0) \right] \\ \left. + \frac{1}{iq_n - p_0 + r_0} \left[ \frac{1}{2} + n_b(p_0) - \frac{1}{2} - n_b(r_0) \right] \right\}$$

Now: \*  $\oint (q_0) = \text{Im } \pi_{\epsilon} \Big|_{iq_n \rightarrow q_0 + i\epsilon}$

\*  $\text{Im } \frac{1}{x \pm i\epsilon} = \mp \pi \delta(x)$

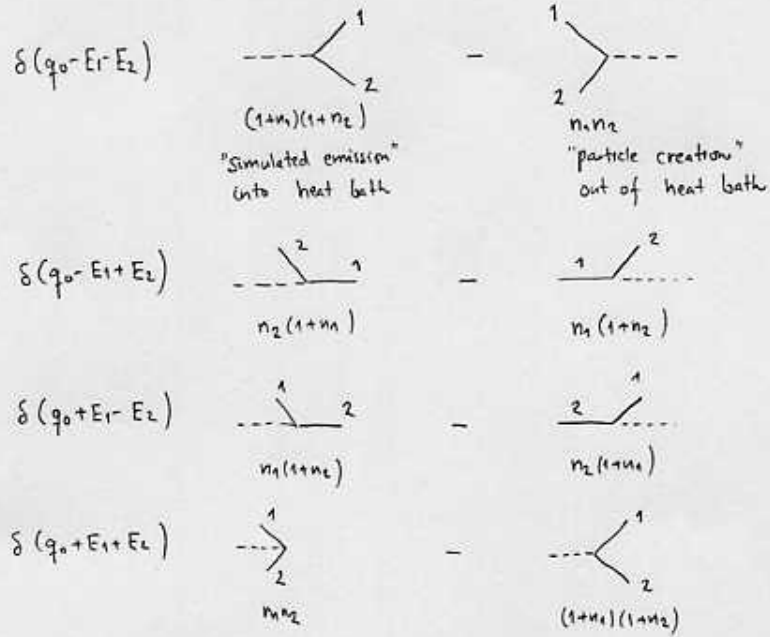
$$\Rightarrow \oint (q_0) = \oint^2 \pi \int \frac{d^3 \vec{p}}{(2\pi)^3} \frac{1}{(z_p)(z_{r_0})} \left\{ \begin{aligned} & \delta(-q_0 - p_0 + r_0) [n_b(p_0) - n_b(r_0)] \\ & - \delta(q_0 + p_0 + r_0) [1 + n_b(p_0) + n_b(r_0)] \\ & + \delta(-q_0 + p_0 + r_0) [1 + n_b(p_0) + n_b(r_0)] \\ & - \delta(q_0 - p_0 + r_0) [n_b(p_0) - n_b(r_0)] \end{aligned} \right\} \quad (*)$$

Or, denoting  $p_0 = E_1$ ,  $n_b(p_0) = n_b(E_1) \equiv n_1$   
 $r_0 = E_2$ ,  $n_b(r_0) = n_b(E_2) \equiv n_2$

$$\oint (q_0) = \oint^2 \pi \int \frac{d^3 \vec{p}}{(2\pi)^3} \frac{1}{(z_{E_1})(z_{E_2})} \left\{ \begin{aligned} & \delta(q_0 + E_1 + E_2) [n_1 n_2 - (1+n_1)(1+n_2)] \\ & + \delta(q_0 + E_1 - E_2) [n_1(1+n_2) - (1+n_1)n_2] \\ & + \delta(q_0 - E_1 + E_2) [n_2(1+n_1) - (1+n_2)n_1] \\ & + \delta(q_0 - E_1 - E_2) [(1+n_1)(1+n_2) - n_1 n_2] \end{aligned} \right\} \quad (**)$$

Physical interpretations

(a) Let us inspect eq. (\*\*) on p. 111. It represents the following processes:



$\Rightarrow S(q_0, \vec{q}) \sim$  disappearance rate of particles with energy  $q_0$ , momentum  $\vec{q}$ ?

Actually it can be shown (see below) that the rate at which  $\phi$ -particles are produced and escape from the plasma, is

$$\frac{d[\Gamma/V]}{d\vec{q}^3} = \frac{1}{(2\pi)^3} \frac{1}{(2q_0)} \Pi^{<}(q_0, \vec{q}) = \frac{1}{(2\pi)^3} \frac{1}{(2q_0)} \cdot 2n_b(q_0) S(q_0, \vec{q})$$

The outcome is that at 1-loop order, the formalism of finite-temperature field theory just reproduces the rate that would be guessed based on classical Boltzmann equations!

$\Rightarrow$  Using  $n_b(q_0) [1 + n_b(p_0) + n_b(r_0)] \delta(q_0 - p_0 - r_0) = n_b(p_0) n_b(r_0) \delta(q_0 - p_0 - r_0)$ ,



$$\frac{\Gamma}{V} = \int \frac{d^3q}{(2\pi)^3 (2q_0)} \frac{d^3p}{(2\pi)^3 (2p_0)} \frac{d^3r}{(2\pi)^3 (2r_0)} (2\pi)^4 \delta^{(4)}(Q - p - r) |g|^2 n_b(p_0) n_b(r_0) (1 + n_b(q_0))$$

+ ...

missing, since the particles produced are assumed to escape!